

$$= - \int \frac{d\Omega \vec{p}^5}{(2\pi)^3} \frac{\partial}{\partial p} \left(\frac{p^4}{E} \right) f_{dm}$$

$$\frac{\partial}{\partial p} \left(\frac{p^4}{E} \right) = \frac{4p^3 E - p^4 \frac{dE}{dp}}{E^2} = \frac{4p^3 E - p^4 \frac{p}{E}}{E^2} = \frac{4p^3 E - p^5}{E^2} = \frac{4p^3}{E} - \frac{p^5}{E^3}$$

$$= - \int \frac{d\Omega \vec{p}^5}{(2\pi)^3} \int dp f_{dm} \frac{\partial}{\partial p} \left(\frac{p^4}{E} \right) = - \int \frac{d\Omega \vec{p}^5}{(2\pi)^3} \int_0^\infty dp f_{dm} \left(\frac{4p^3}{E} - \frac{p^5}{E^3} \right) \approx 0 \left(\frac{p^3}{E} \right) \sim v^3!!!$$

$$\approx -4 \int \frac{d\Omega \vec{p}^5}{(2\pi)^3} \int_0^\infty dp \frac{p^3}{E} f_{dm} = -4 \int \frac{d\Omega}{(2\pi)^3} \int_0^\infty dp p^2 \frac{p \vec{p}^5}{E} f_{dm} = -4 \int \frac{d^3 p}{(2\pi)^3} \frac{p \vec{p}^5}{E} f_{dm} =$$

$$= -4 n_{dm} v_J$$

$$\rightarrow \int \frac{d^3 p}{(2\pi)^3} \frac{\partial f_{dm}}{\partial E} \frac{p \vec{p}^5 p^2}{E} = \int \frac{d^3 p}{(2\pi)^3} \frac{E}{p} \frac{\partial f_{dm}}{\partial p} \frac{p \vec{p}^5 p^2}{E} = \int \frac{d^3 p}{(2\pi)^3} p \vec{p}^5 \frac{\partial f_{dm}}{\partial p} =$$

$$= \int d\Omega \vec{p}^5 \int \frac{dp}{(2\pi)^3} p^3 \frac{\partial f_{dm}}{\partial p} = -\delta^{i5} \frac{4\pi}{3} \int \frac{dp}{8\pi^3} 3p^2 f_{dm} = -\delta^{i5} \frac{4\pi}{8\pi^3} \int dp p^2 f_{dm} =$$

$$= -\delta^{i5} \frac{4\pi}{3} \int dp p^2 f_{dm} = -\delta^{i5} \frac{4\pi}{3} \int d^3 p f_{dm} = \frac{1}{2\pi^2} \int d^3 p f_{dm} = n_{dm}$$

$$= -\delta^{i5} n_{dm}$$

$$\rightarrow -\frac{1}{a} \frac{\partial \Psi}{\partial x^i} \int \frac{d^3 p}{(2\pi)^3} \frac{\partial f_{dm}}{\partial E} \frac{p \vec{p}^5 p^2}{E} = -\frac{1}{a} \frac{\partial \Psi}{\partial x^i} \times -\delta^{i5} n_{dm} = \frac{n_{dm}}{a} \frac{\partial \Psi}{\partial x^5}$$

Putting everything together 1st moment of Boltzmann equation:

$$\frac{\partial (n_{dm} v_J)}{\partial t} - 4 n_{dm} v_J \times \left[\frac{da/dt}{a} + \frac{\partial \vec{\sigma}}{\partial t} \right] + \frac{n_{dm}}{a} \frac{\partial \Psi}{\partial x^5} \approx \frac{\partial (n_{dm} v_J)}{\partial t} + 4 H n_{dm} v_J + \frac{n_{dm}}{a} \frac{\partial \Psi}{\partial x^5} = 0$$

n, Ψ 1st order, so there are no 0th order terms

\rightarrow can set $n_{dm} = n_{dm}^{(0)}$ everywhere (otherwise we get 2nd order terms)

$$\frac{\partial}{\partial t} (n_{dm} v^j) + 4H n_{dm} v^j + \frac{n_{dm}}{a} \frac{\partial \Psi}{\partial x^j} = 0$$

↓

$$\frac{\partial}{\partial t} (n_{dm}^{(0)} v^j) + 4H n_{dm}^{(0)} v^j + \frac{n_{dm}^{(0)}}{a} \frac{\partial \Psi}{\partial x^j} = 0$$

$$\downarrow n_{dm}^{(0)} \propto a^{-3} \rightarrow \frac{\partial}{\partial t} n_{dm}^{(0)} = -3H n_{dm}^{(0)}$$

$$-3H n_{dm}^{(0)} v^j + n_{dm}^{(0)} \frac{\partial v^j}{\partial t} + 4H n_{dm}^{(0)} v^j + \frac{n_{dm}^{(0)}}{a} \frac{\partial \Psi}{\partial x^j} = 0$$

⇓

$$\boxed{\frac{\partial v^j}{\partial t} + H v^j + \frac{1}{a} \frac{\partial \Psi}{\partial x^j} = 0}$$

Boltzmann equation for cold DM velocity (1st order / 1st moment)

↳ no $(\vec{v} \cdot \nabla) \vec{v}$ term since it is 2nd order!

Note: the equation for the n th moment depends on the $(n+1)$ th moment → infinite hierarchy (in principle)

- ① $\delta \rightarrow v$
 - ② $\sigma \rightarrow$ quadrupole/anisotropic stress?
 - ③ sufficient to characterize CDM!!!
- ↳ generic feature of integrating Boltzmann equations to get fluid equations

Dropped terms $O((\frac{v}{c})^2)$!!!

Because DM is cold

Conformal time

Equations for δ and v

$$\frac{\partial \delta}{\partial t} + \frac{1}{a} \frac{\partial v^i}{\partial x^i} + 3\frac{\partial \Phi}{\partial t} = 0$$

$$d\eta = \frac{dt}{a} \rightarrow \frac{d}{dt} = \frac{1}{a} \frac{d}{d\eta} = \frac{1}{a} \frac{d}{dt}$$

$$\delta + \frac{\partial v^i}{\partial x^i} + 3\dot{\Phi} = 0$$

$$\frac{\partial v^j}{\partial t} + H v^j + \frac{1}{a} \frac{\partial \Psi}{\partial x^j} = 0$$

$$\dot{v}_j + \underbrace{\frac{\dot{a}}{a} v_j}_{H v_j} + \frac{\partial \Psi}{\partial x^j} = 0$$

As usual take Fourier transform, assume irrotational

velocity $(\vec{v} \propto \vec{k} \rightarrow \tilde{v}^i = \left(\frac{k^i}{k}\right) \tilde{v} \rightarrow \left(\frac{\partial v}{\partial x^i}\right) = ik \tilde{v}$)

$$\begin{cases} \dot{\delta} + ik \tilde{v} + 3 \dot{\Phi} = 0 \\ \ddot{\tilde{v}} + \frac{\dot{a}}{a} \tilde{v} + ik \tilde{\Psi} = 0 \end{cases}$$

Collisionless Boltzmann equation for DRs

- to 1st order
- in Fourier space
- expressed in terms of 2 lowest order moments δ, v (since DR is cold!!!)

Boltzmann equations for baryons

Baryons: electrons + protons (bad nomenclature)

coupled by Coulomb scattering $e+p \rightarrow e+p$
rate always larger than H.

$$\frac{\rho_e - \rho_e^{(0)}}{\rho_e^{(0)}} = \frac{\rho_p - \rho_p^{(0)}}{\rho_p^{(0)}} \equiv \delta_b$$

strong scattering rate forces common overdensity

$$\bar{v}_e = \bar{v}_p \equiv \bar{v}_b$$

similarly for velocity

So similarly to ~~dark~~ dark matter we want the equations for δ_b, v_b

Start from unintegrated equations for e, p :

$$\frac{d\rho_e}{dt}(\vec{x}, \vec{q}, t) = \langle C_{ep} \rangle_{qq'q'} + \langle C_{ex} \rangle_{pp'q'} \rightarrow \text{Compton scattering}$$

$e(\vec{q}) + \gamma(\vec{p}) \rightarrow e(\vec{q}') + \gamma(\vec{p}')$

$$\frac{d\rho_p}{dt}(\vec{x}, \vec{Q}, t) = \langle C_{ep} \rangle_{qq'Q}$$

Coulomb scattering
 $e(\vec{q}) + p^+(\vec{Q}) \rightarrow e^-(\vec{q}') + p^+(\vec{Q}')$

Compact notation (compare Boltzmann equations for photons)

Unintegrated part of collision term $\rightarrow z^3$

$$C_{e\gamma} \equiv (2\pi)^4 \delta^4(p+q-p'-q') \frac{|M|_{\text{Compton}}^2}{\delta E(p)E(p')E_e(q')E_e(q)} [f_e(q')f_\gamma(p') - f_e(q)f_\gamma(p)]$$

$$\langle \dots \rangle_{pp'q'} \equiv \int \frac{d^3p}{(2\pi)^3} \int \frac{d^3p'}{(2\pi)^3} \int \frac{d^3q'}{(2\pi)^3} (\dots)$$

Similarly for $C_{ep} \approx |M|_{\text{Compton}}^2 \rightarrow |M|_{\text{Coulomb}}^2$ and integrated

momenta are different

$\langle C_{ep} \rangle$ neglected since $\sigma \propto \frac{1}{m_p^2}$, so interactions of combined electron-proton fluid with photons driven by Compton scattering between electrons and photons

(Also neglecting ionization and recombination terms)

Now multiply $\frac{df_e}{dt}$ equation by $\frac{d^3q}{(2\pi)^3}$ and integrate \rightarrow LHS identical to n_e , recall:

$$\frac{dn_e}{dt} = \frac{\partial f_e}{\partial t} + \frac{\partial f_e}{\partial x^i} \frac{p^i}{a} \frac{1}{E} - \frac{\partial f_e}{\partial E} \left[H \frac{p^2}{E} + \frac{p^2}{E} \frac{\partial \Phi}{\partial t} + \frac{p^i}{a} \frac{\partial \Psi}{\partial x^i} \right] = 0$$

$$\downarrow \int \times \frac{d^3q}{(2\pi)^3}$$

~~$$\frac{\partial n_e}{\partial t} + \frac{1}{a} \frac{\partial (n_e v_{ei}^i)}{\partial x^i} + 3 \left[H + \frac{\partial \Phi}{\partial t} \right] n_e = \langle C_{ep} \rangle_{\partial q q' q} + \langle C_{e\gamma} \rangle_{pp' q' q}$$~~

$$\frac{\partial n_e}{\partial t} + \frac{1}{a} \frac{\partial (n_e v_{ei}^i)}{\partial x^i} + 3 \left[H + \frac{\partial \Phi}{\partial t} \right] n_e = \langle C_{ep} \rangle_{\partial q q' q} + \langle C_{e\gamma} \rangle_{pp' q' q} \rightarrow 0$$

$\equiv 0$ mathematical physical identity, not approximation

Why? Take e.g.

$$\langle C_{ep} \rangle_{\alpha\alpha'q'q}$$

$\langle \dots \rangle_{\alpha\alpha'q'q}$ symmetric in $\alpha \leftrightarrow \alpha'$ $q \leftrightarrow q'$

C_{ep} antisymmetric in $\alpha \leftrightarrow \alpha'$ $q \leftrightarrow q'$ because of

~~$$f_e(q')f_p(\alpha) - f_e(q)f_p(\alpha')$$~~

$$f_e(q)f_p(\alpha) - f_e(q')f_p(\alpha')$$

So integral is identically 0!

Physically speaking, Coulomb scattering conserves electron number,

so cannot contribute to $\frac{\partial n_e}{\partial t}$

$$\int f_e(q')f_p(\alpha') \propto \# \text{ electrons produced in Coulomb scattering}$$

$$\int f_e(q)f_p(\alpha) \propto \# \text{ electrons lost in Coulomb scattering}$$

Similarly for Compton scattering

True in general: multiplying unintegrated collision term by

conserved quantity and integrating $\rightarrow 0$

So for S_b we have the exact same equation as cold DR.

Using conformal time and going to Fourier space:

$$\boxed{\dot{S}_b + ik \tilde{v}_b + 3\tilde{\Phi} = 0}$$

Boltzmann equation for baryon overdensity (1st order/0th moment)

As with DR we also need an equation for the velocity, take first

moments of $\frac{df_e}{dt}$ and $\frac{df_p}{dt}$ equations

~~$$\int \frac{df_e}{dt} \frac{d^3q}{(2\pi)^3} \frac{d^3\alpha}{E}$$~~
~~$$\int \frac{df_p}{dt} \frac{d^3q}{(2\pi)^3} \frac{d^3\alpha}{E}$$~~

$$\int \frac{d^3q}{(2\pi)^3} \bar{q} \frac{d\mathcal{L}_2}{dt}(\bar{x}, \bar{q}, t) = \langle C_{ep} \bar{q} \rangle_{qq'q''} + \langle C_{er} \bar{q} \rangle_{pp'q''}$$

$$\int \frac{d^3Q}{(2\pi)^3} \bar{Q} \frac{d\mathcal{L}_2}{dt}(\bar{x}, \bar{Q}, t) = \langle C_{ep} \bar{Q} \rangle_{qq'Q''}$$

Recall for m we did the same but with $\int \frac{d^3p}{(2\pi)^3} \frac{p p^i}{E} = \int \frac{d^3p}{(2\pi)^3} \frac{\vec{p}}{E}$

So here we can borrow the same results up to a factor of E_{m_e, m_p}

For m we got

$$\frac{\partial(n_{dm} v^j)}{\partial t} + 4H n_{dm} v^j + \frac{n_{dm}}{a} \frac{\partial \Psi}{\partial x^j} = 0$$

Here we get:

$$m_e \frac{\partial(n_b v_b^j)}{\partial t} + 4H n_b v_b^j + \frac{n_b m_e}{a} \frac{\partial \Psi}{\partial x^j} = \langle C_{ep} \bar{q} \rangle_{qq'q''} + \langle C_{er} \bar{q} \rangle_{pp'q''}$$

$$+ m_p \frac{\partial(n_b v_b^j)}{\partial t} + 4H m_p n_b v_b^j + \frac{n_b m_p}{a} \frac{\partial \Psi}{\partial x^j} = \langle C_{ep} \bar{Q} \rangle_{qq'Q''}$$

$$(m_e + m_p) \left[\frac{\partial(n_b v_b^j)}{\partial t} + 4H n_b v_b^j + \frac{n_b}{a} \frac{\partial \Psi}{\partial x^j} \right] = \langle C_{ep} (q^j + Q^j) \rangle_{qq'q''} + \langle C_{er} q^j \rangle_{pp'q''}$$

$m_p \gg m_e$

otherwise get
 2nd order terms

again mathematically
 since $q^j + Q^j$ symmetric
AND conserved!!

$$n_b \rightarrow n_b^{(0)} \propto a^{-3} \quad \text{so} \quad \frac{\partial n_b^{(0)}}{\partial t} = -3H n_b^{(0)}$$

$$m_p \frac{\partial n_b^{(0)} v_b^j}{\partial t} + m_p n_b^{(0)} \frac{\partial v_b^j}{\partial t} + 4H m_p n_b^{(0)} v_b^j + \frac{m_p n_b^{(0)}}{a} \frac{\partial \Psi}{\partial x^j} = \langle C_{er} q^j \rangle_{pp'q''}$$

$$-3H m_p n_b^{(0)} v_b^j + m_p n_b^{(0)} \frac{\partial v_b^j}{\partial t} + 4H m_p n_b^{(0)} v_b^j + \frac{m_p n_b^{(0)}}{a} \frac{\partial \Psi}{\partial x^j} = \langle C_{er} q^j \rangle_{pp'q''}$$

$$\rightarrow \frac{n_p n_b^{(0)} \frac{\partial v_b^j}{\partial t}}{\rho_b} + \frac{H n_p n_b^{(0)} v_b^j}{\rho_b} + \frac{n_p n_b^{(0)} \frac{\partial \psi}{\partial x^j}}{a \rho_b} = \frac{\langle C_{ex} q^j \rangle_{pp'q'q}}{\rho_b}$$

$$\frac{\partial v_b^j}{\partial t} + H v_b^j + \frac{1}{a} \frac{\partial \psi}{\partial x^j} = \frac{1}{\rho_b} \langle C_{ex} q^j \rangle_{pp'q'q}$$

$\langle C_{ex} \bar{q} \rangle_{pp'q'q} = - \langle C_{ex} \bar{p} \rangle_{pp'q'q}$ since $\langle C_{ex} (\bar{q} + \bar{p}) \rangle_{pp'q'q} = 0$
 by the conservation of momentum

• $-\frac{\langle C_{ex} \bar{p} \rangle_{pp'q'q}}{\rho_b} \rightarrow$ go to Fourier space
 \rightarrow multiply by \hat{k}_j (similarly for LHS)

$$\hat{k} \cdot \bar{p} = p_\mu \rightarrow - \frac{\langle C_{ex} p_\mu \rangle_{pp'q'q}}{\rho_b}$$

Recall we already ρ_b computed $\langle C_{ex} \rangle_{pp'q'q}$ for the photon Boltzmann equation

$$\langle C_{ex} \rangle_{pp'q'q} = -p \frac{\partial f^{(0)}}{\partial p} n_e \sigma_T [\Theta_0 - \Theta(\hat{p}) + \hat{p} \cdot \vec{v}_b]$$

$$-\frac{\langle C_{ex} p_\mu \rangle_{pp'q'q}}{\rho_b} = \frac{n_e \sigma_T}{\rho_b} \int \frac{d^3 p}{(2\pi)^3} p^2 \frac{\partial f^{(0)}}{\partial p} \mu [\tilde{\Theta}_0 - \tilde{\Theta}(\mu) + \hat{p} \cdot \vec{v}_b] =$$

$$= \frac{n_e \sigma_T}{\rho_b} \int_0^\infty \frac{dp}{2\pi^2} p^4 \frac{\partial f^{(0)}}{\partial p} \int_{-1}^1 \frac{d\mu}{2} \mu [\tilde{\Theta}_0 - \tilde{\Theta}(\mu) + \hat{p} \cdot \vec{v}_b]$$

$\int \frac{d^3 p}{(2\pi)^3} = \int \frac{dp p^2}{(2\pi)^3} \int d\varphi \int d(\cos\theta) = \int \frac{dp p^2}{8\pi^3} 2\pi \int d\mu = \int \frac{dp p^2}{4\pi^2} \int d\mu = \int \frac{dp p^2}{2\pi^2} \int \frac{d\mu}{2}$

$$\frac{n_e \sigma_T}{b} \int_0^\infty \frac{dp}{2\pi^2} p^4 \frac{\partial f^{(e)}}{\partial p} \int_{-1}^1 \frac{d\mu}{2} \mu [\tilde{\Theta}_0 - \tilde{\Theta}(\mu) + \tilde{v}_b \mu]$$

$$\Rightarrow \int_0^\infty \frac{dp}{2\pi^2} p^4 \frac{\partial f^{(e)}}{\partial p} = -4 \int_0^\infty \frac{dp}{2\pi^2} p^3 f^{(e)} = -4 \int_0^\infty \frac{dp p^2}{2\pi^2} p f^{(e)} = -4 \int \frac{d^3p}{(2\pi)^3} p f^{(e)} = -4\rho_p$$

$$\int_{-1}^1 \frac{d\mu}{2} \mu [\tilde{\Theta}_0 - \tilde{\Theta}(\mu) + \tilde{v}_b \mu] = i\tilde{\Theta}_1 + \frac{\tilde{v}_b}{3}$$

0 (odd)

$$\int_{-1}^1 \frac{d\mu}{2} \tilde{v}_b \mu^2 = \tilde{v}_b \frac{\mu^3}{6} \Big|_{-1}^1 = \tilde{v}_b \left(\frac{1}{6} + \frac{1}{6} \right) = \frac{\tilde{v}_b}{3}$$

Recall

$$\Theta_\ell \equiv \frac{1}{(-i)^\ell} \int_{-1}^1 \frac{d\mu}{2} P_\ell(\mu) \Theta(\mu) \quad P_1(\mu) = \mu$$

$$\text{so } i \int_{-1}^1 \frac{d\mu}{2} \mu \tilde{\Theta}(\mu) = \tilde{\Theta}_1 \rightarrow - \int \frac{d\mu}{2} \mu \tilde{\Theta}(\mu) = i\tilde{\Theta}_1$$

$$\rightarrow = \frac{n_e \sigma_T}{4 b} \rho_p \left(i\tilde{\Theta}_1 + \frac{\tilde{v}_b}{3} \right)$$

~~Putting everything together~~

~~$$\frac{\partial v_b^{\text{J}}}{\partial t} + \frac{da/dt}{a} v_b^{\text{J}} + \frac{1}{a} \frac{\partial \psi}{\partial x^{\text{J}}} = - \frac{n_e \sigma_T}{b} \rho_p \left(i\tilde{\Theta}_1 + \frac{\tilde{v}_b}{3} \right)$$~~

Putting everything together

$$\text{LHS: } \frac{\partial v_b^{\text{J}}}{\partial t} + \frac{da/dt}{a} v_b^{\text{J}} + \frac{1}{a} \frac{\partial \psi}{\partial x^{\text{J}}} \xrightarrow{t \rightarrow \eta} \frac{1}{a} \dot{v}_b + \frac{\dot{a}}{a^2} v_b^{\text{J}} + \frac{1}{a} \frac{\partial \psi}{\partial x^{\text{J}}} \xrightarrow{\text{FT}} \frac{1}{a} \dot{\tilde{v}}_b + \frac{\dot{a}}{a^2} \tilde{v}_b + \frac{i k \tilde{\psi}}{a}$$

$$\text{RHS: } -4 \frac{n_e \sigma_T}{b} \rho_p \left(i\tilde{\Theta}_1 + \frac{\tilde{v}_b}{3} \right) \quad \text{already in Fourier space!}$$

$$\rightarrow \frac{1}{a} \dot{\tilde{v}}_b + \frac{\dot{a}}{a^2} \tilde{v}_b + \frac{i k \tilde{\psi}}{a} = -4 \frac{n_e \sigma_T}{b} \rho_p \left(i\tilde{\Theta}_1 + \frac{\tilde{v}_b}{3} \right)$$

$$\Rightarrow \dot{\tilde{v}}_b + \frac{\dot{a}}{a} \tilde{v}_b + ik\tilde{\psi} = -4n_e \sigma_T a \frac{\rho_\gamma}{\rho_b} (i\tilde{\theta}_1 + \frac{\tilde{v}_b}{3})$$

Recall definition of optical depth

$$\tau(\eta) \equiv \int_{\eta}^{\eta_0} d\eta' n_e \sigma_T a \rightarrow \dot{\tau} \equiv \frac{d\tau}{d\eta} = -n_e \sigma_T a$$

$$\boxed{\dot{\tilde{v}}_b + \frac{\dot{a}}{a} \tilde{v}_b + ik\tilde{\psi} = \dot{\tau} \frac{4\rho_\gamma}{3\rho_b} [3i\tilde{\theta}_1 + \tilde{v}_b]}$$

Boltzmann equation for baryon velocity (1st order/1st moment)

$$\tilde{\theta}_1 \sim \tilde{v}_\gamma \quad \text{so} \quad [\dots] \sim v_\gamma - v_b \quad \text{: momentum exchange}$$

Physical meaning of $\frac{4}{3}$ factor - moving e^- is difficult because they are tightly coupled to heavy protons. If $m_p \rightarrow \infty$, $\rho_b \rightarrow \infty$, Compton scattering cannot change electron velocity

$$R \equiv \frac{3\rho_b}{4\rho_\gamma} = \frac{(1+w_b)\rho_b}{(1+w_\gamma)\rho_\gamma} \quad \text{ph baryon-to-photon density ratio}$$

* even if there is neutral hydrogen so $n_p \neq n_b$, approximately valid

Recap Boltzmann equations so far

$$\dot{\tilde{\theta}} + ik_\mu \tilde{\theta} + \dot{\tilde{\Phi}} + ik_\mu \tilde{\psi} = -\dot{\tau} [\tilde{\theta}_0 - \tilde{\theta} + \mu \tilde{v}_b]$$

$$\dot{\tilde{\delta}}_{dm} + ik \tilde{v}_{dm} = -3\dot{\tilde{\Phi}}$$

$$\dot{\tilde{v}}_{dm} + \frac{\dot{a}}{a} \tilde{v}_{dm} = -ik\tilde{\psi}$$

$$\dot{\tilde{\delta}}_b + ik\tilde{v}_b = -3\dot{\tilde{\Phi}}$$

$$\dot{\tilde{v}}_b + \frac{\dot{a}}{a} \tilde{v}_b = -ik\tilde{\psi} + \frac{\dot{\tau}}{R} [\tilde{v}_b + 3i\tilde{\theta}_1]$$

Summary of Boltzmann equations for γ, DM, b, ν

+ other effects we glossed over
(Drop \sim from Fourier transforms)

↑
massless ν basically like γ , just not interacting! \rightarrow no scattering

$$\dot{\Theta} + ik_{\mu} \Theta = -\dot{\Phi} - ik_{\mu} \Psi - \dot{\tau} \left[\Theta_0 - \Theta + \mu v_b - \frac{1}{2} P_2(\mu) \Pi \right]$$

$\Pi = \Theta_2 + \Theta_{p2} + \Theta_{p0}$ } temperature field also coupled + polarization field

$$\dot{\Theta}_p + ik_{\mu} \Theta_p = -\dot{\tau} \left[-\Theta_p + \frac{1}{2} (1 - P_2(\mu)) \Pi \right]$$

$$\dot{\delta}_{dm} + ik_{\mu} v_{dm} = -3\dot{\Phi}$$

$$\dot{v} + \frac{\dot{a}}{a} v = -ik\Psi$$

$$\dot{\delta}_b + ik v_b = -3\dot{\Phi}$$

$$\dot{\delta}_b + \frac{\dot{a}}{a} v_b = -ik\Psi + \frac{\dot{\tau}}{R} [v_b + 3i\Theta_1]$$

$$\dot{N} + ik_{\mu} N = -\dot{\Phi} - ik_{\mu} \Psi$$

angular dependence of Compton scattering
 $\mu^2 \propto |\hat{\epsilon} \cdot \hat{\epsilon}'|^2$
evolution of polarization field

These equations tell us how particle distributions are affected by metric fluctuations. Now we need to know the other way around \rightarrow equations for $\Phi, \Psi!$
 \downarrow
(perturbed) Einstein equations

massless neutrinos, without scattering, so equation identical to that for photons, minus scattering

\rightarrow typically expanded in multipoles

$$\Theta_{\ell} \equiv \frac{1}{(-i)^{\ell}} \int_{-1}^1 d\mu \frac{d^{\ell}}{d\mu^{\ell}} P_{\ell}(\mu) \Theta(\mu)$$

e.g. first two moments neglecting scattering

$$\begin{cases} \dot{\Theta}_0 + k \Theta_1 = -\dot{\Phi} \\ \dot{\Theta}_1 - \frac{k}{3} \Theta_0 = -\frac{k}{3} \dot{\Phi} \end{cases}$$